THE INVERSE MEAN CURVATURE FLOW IN
COSMOLOGICAL SPACETIMES

CLAUS GERHARDT

Abstract. We prove that the leaves of an inverse mean curvature flow provide a foliation of a future end of a cosmological spacetime $N$ under the necessary and sufficient assumptions that $N$ satisfies a future mean curvature barrier condition and a strong volume decay condition. Moreover, the flow parameter $t$ can be used to define a new physically important time function.

Contents

0. Introduction 1
1. Notations and definitions 4
2. The evolution problem 9
3. Lower order estimates 10
4. $C^1$-estimates 14
5. $C^2$-estimates 17
6. Longtime existence 19
7. A new time function 19
References 21

0. Introduction

The inverse mean curvature flow has already been considered in Euclidean space [3] or in asymptotically flat Riemannian spaces [12]. In the latter case Huisken and Ilmanen used it to prove the Penrose inequality. One major difficulty in their proof was that jumps might occur during the flow, i.e., the mean curvature of the flow hypersurfaces might vanish even though the initial hypersurface has positive mean curvature.

The Lorentzian geometry is much more favourable for curvature flows, cf. [1] [5] [6] [8], so that no jumps should occur in case of the inverse mean

Date: 5th March 2004.
2000 Mathematics Subject Classification. 35J60, 53C21, 53C44, 53C50, 58J05.
Key words and phrases. Lorentzian manifold, cosmological spacetime, general relativity, inverse mean curvature flow.

This work has been supported by the Deutsche Forschungsgemeinschaft.
curvature flow. We shall show that this is indeed the case, if the ambient space is a globally hyperbolic \((n+1)\)-dimensional Lorentzian manifold \(N\) with a compact Cauchy hypersurface satisfying the timelike convergence condition

\[ \bar{R}_{\alpha\beta} \nu^\alpha \nu^\beta \geq 0 \quad \forall \langle \nu, \nu \rangle = -1. \]  

Such spaces are called cosmological spacetimes, a terminology due to Bartnik.

Let \(M_0 \subset N\) be a spacelike hypersurface the mean curvature of which is either strictly positive or negative, then we consider the inverse mean curvature flow (IMCF)

\[ \dot{x} = -H^{-1} \nu \]

with initial hypersurface \(M_0\). Here, \(\nu\) is the past directed normal of the flow hypersurfaces \(M(t)\) and \(H = H_{|M(t)}\) the corresponding mean curvature, i.e., the trace of the second fundamental form.

If \(H_{|M_0}\) is positive resp. negative, then the flow moves to the future resp. past of \(M_0\). Furthermore, \(H_{|M(t)}\) will uniformly tend to \(\infty\) resp. \(-\infty\), if the flow exists for all time.

In former papers we referred to this latter phenomenon by saying that there were crushing singularities in the future resp. past, erroneously assuming that only big crunch or big bang type singularities could produce spacelike hypersurfaces the mean curvatures of which become unbounded if the hypersurfaces approached the singularities.

But a behaviour like that could also be caused by a null hypersurface \(\mathcal{H}\), e.g., by the event horizon of a black hole, if the spacetime can be viewed as having a past or future boundary component \(\mathcal{H}\) that can be identified with a compact null hypersurface representing a non-crushing singularity, i.e., the Riemannian curvature tensors remains uniformly bounded near \(\mathcal{H}\)

\[ \bar{R}_{\alpha\beta\gamma\delta} \bar{R}^{\alpha\beta\gamma\delta} \leq \text{const}. \]

An example of such a spacetime is given in Section 4.

We therefore define

0.1. Definition. Let \(N\) be a globally hyperbolic spacetime with compact Cauchy hypersurface \(\mathcal{S}_0\) so that \(N\) can be written as a topological product \(N = \mathbb{R} \times \mathcal{S}_0\) and its metric expressed as

\[ ds^2 = e^{2\varphi}(-(dx^0)^2 + \sigma_{ij}(x^0, x)dx^i dx^j). \]

Here, \(x^0\) is a globally defined future directed time function and \((x^i)\) are local coordinates for \(\mathcal{S}_0\). \(N\) is said to have a future mean curvature barrier resp. past mean curvature barrier, if there are sequences \(M_k^+\) resp. \(M_k^-\) of closed spacelike hypersurfaces such that

\[ \lim_{k \to \infty} H_{|M_k^+} = \infty \quad \text{resp.} \quad \lim_{k \to \infty} H_{|M_k^-} = -\infty. \]
and
\[(0.6) \quad \limsup_{M^+_x} \inf_{x^0} > x^0(p) \quad \forall \ p \in N \]
resp.
\[(0.7) \quad \liminf_{M^-_x} \sup_{x^0} < x^0(p) \quad \forall \ p \in N. \]

0.2. Remark. Let $N$ be a cosmological spacetime with future and a past mean curvature barriers, then it can be foliated by closed hypersurfaces of constant mean curvature, cf. [2]. Moreover, the mean curvature function $\tau$ is continuous in $N$ and smooth in $\{\tau \neq 0\}$ with non-vanishing gradient, hence it can be used as a time function, cf. [9]. These results are also valid in future resp. past ends.

We shall assume in the following that $N$ has a future mean curvature barrier. By reversing the time direction this configuration also comprises the case that $N$ has a past mean curvature barrier.

Under this assumption we shall prove that, for a given compact spacelike hypersurface $M_0$ with $H_{|M_0} > 0$, the future of $M_0$ can be foliated by the leaves of an IMCF starting at $M_0$ provided a so-called future strong volume decay condition is satisfied, cf. Definition 1.2. A strong volume decay condition is both necessary and sufficient in order that the IMCF exists for all time.

The main result of this paper can be summarized in the following theorem

0.3. Theorem. Let $N$ be a cosmological spacetime with compact Cauchy hypersurface $S_0$ and with a future mean curvature barrier. Let $M_0$ be a closed spacelike hypersurface with positive mean curvature and assume furthermore that $N$ satisfies a future volume decay condition. Then the IMCF 0.2 with initial hypersurface $M_0$ exists for all time and provides a foliation of the future $D_+(M_0)$ of $M_0$.

The evolution parameter $t$ can be chosen as a new time function. The flow hypersurfaces $M(t)$ are the slices $\{t = \text{const}\}$ and their volume satisfies
\[(0.8) \quad |M(t)| = |M_0|e^{-t}.\]

Defining an almost proper time function $\tau$ by choosing
\[(0.9) \quad \tau = 1 - e^{-\frac{1}{2}t}\]
we obtain $0 \leq \tau < 1$,
\[(0.10) \quad |M(\tau)| = |M_0|(1 - \tau)^n, \]
and the future singularity corresponds to $\tau = 1$.

Moreover, the length $L(\gamma)$ of any future directed curve $\gamma$ starting from $M(\tau)$ is bounded from above by
\[(0.11) \quad L(\gamma) \leq c(1 - \tau), \]
where \( c = c(n, M_0) \). Thus, the expression \( 1 - \tau \) can be looked at as the radius of the slices \( \{ \tau = \text{const} \} \) as well as a measure of the remaining life span of the universe.

Without any further structural assumptions it seems impossible to derive any convergence results for an appropriately rescaled IMCF. In [10] we look at the IMCF in asymptotically Robertson Walker spaces and prove that a properly rescaled flow converges indeed.

1. Notations and definitions

The main objective of this section is to state the equations of Gauß, Codazzi, and Weingarten for spacelike hypersurfaces \( M \) in a \((n+1)\)-dimensional Lorentzian manifold \( N \). Geometric quantities in \( N \) will be denoted by \((\bar{g}_{\alpha\beta})\), \((\bar{R}_{\alpha\beta\gamma\delta})\), etc., and those in \( M \) by \((g_{ij})\), \((R_{ijkl})\), etc. Greek indices range from 0 to \( n \) and Latin from 1 to \( n \); the summation convention is always used. Generic coordinate systems in \( N \) resp. \( M \) will be denoted by \((x^\alpha)\) resp. \((\xi^i)\). Covariant differentiation will simply be indicated by indices, only in case of possible ambiguity they will be preceded by a semicolon, i.e., for a function \( u \) in \( N \), \((u_\alpha)\) will be the gradient and \((u_{\alpha\beta})\) the Hessian, but e.g., the covariant derivative of the curvature tensor will be abbreviated by \( \bar{R}_{\alpha\beta\gamma\delta,\epsilon} \).

We also point out that

\[
\bar{R}_{\alpha\beta\gamma\delta,\epsilon} = \bar{R}_{\alpha\beta\gamma\delta x^i} \xi^i
\]

with obvious generalizations to other quantities.

Let \( M \) be a spacelike hypersurface, i.e., the induced metric is Riemannian, with a differentiable normal \( \nu \) which is timelike.

In local coordinates, \((x^\alpha)\) and \((\xi^i)\), the geometric quantities of the spacelike hypersurface \( M \) are connected through the following equations

1. 

\[
x^\alpha_{ij} = h_{ij} \nu^\alpha
\]

the so-called Gauß formula. Here, and also in the sequel, a covariant derivative is always a full tensor, i.e.

2. 

\[
x^\alpha_{ij} = x^\alpha_{,ij} - \Gamma^k_{ij} x^\alpha_k + \bar{\Gamma}^\alpha_{\beta\gamma} x^\beta_i x^\gamma_j
\]

The comma indicates ordinary partial derivatives. In this implicit definition the second fundamental form \((h_{ij})\) is taken with respect to \( \nu \).

The second equation is the Weingarten equation

3. 

\[
\nu^\alpha_{ij} = h^k_{ij} \nu^\alpha_k
\]

where we remember that \( \nu^\alpha_{ij} \) is a full tensor. Finally, we have the Codazzi equation

4. 

\[
h_{ij,k} - h_{ik,j} = \bar{R}_{\alpha\beta\gamma\delta} \nu^\alpha_{ij} x^\gamma_j x^\delta_k
\]
and the *Gauß equation*

\[ R_{ijkl} = -\{ h_{ik}h_{jl} - h_{il}h_{jk} \} + \bar{R}_{\alpha\beta\gamma\delta}x^\alpha_i x^\beta_j x^\gamma_k x^\delta_l. \]

Now, let us assume that \( N \) is a globally hyperbolic Lorentzian manifold with a *compact* Cauchy surface. \( N \) is then a topological product \( I \times S_0 \), where \( I \) is an open interval, \( S_0 \) is a compact Riemannian manifold, and there exists a Gaussian coordinate system \( (x^\alpha) \), such that the metric in \( N \) has the form

\[ ds_N^2 = e^{2\psi}\{-dx^0^2 + \sigma_{ij}(x^0, x)dx^i dx^j\}, \]

where \( \sigma_{ij} \) is a Riemannian metric, \( \psi \) a function on \( N \), and \( x \) an abbreviation for the spacelike components \( (x^i) \). We also assume that the coordinate system is future oriented, i.e., the time coordinate \( x^0 \) increases on future directed curves. Hence, the contravariant timelike vector \( (\xi^\alpha) = (1, -1, 0, \ldots, 0) \) is future directed as is its covariant version \( (\xi^\alpha) = e^{-2\psi}(1, u^i) \).

Let \( M = \text{graph } u\big|_{S_0} \) be a spacelike hypersurface

\[ M = \{ (x^0, x): x^0 = u(x), x \in S_0 \}, \]

then the induced metric has the form

\[ g_{ij} = e^{2\psi}\{-u_i u_j + \sigma_{ij}\} \]

where \( \sigma_{ij} \) is evaluated at \((u, x)\), and its inverse \((g^{ij}) = (g_{ij})^{-1}\) can be expressed as

\[ g^{ij} = e^{-2\psi}\{\sigma^{ij} + u_i u_j/v^2\}, \]

where \((\sigma^{ij}) = (\sigma_{ij})^{-1}\) and

\[ u^i = \sigma^{ij} u_j \]
\[ v^2 = 1 - \sigma^{ij} u_i u_j \equiv 1 - |Du|^2. \]

Hence, graph \( u \) is spacelike if and only if \(|Du| < 1\).

The covariant form of a normal vector of a graph looks like

\[ (\nu^\alpha) = \pm v^{-1} e^\psi(1, -u_i). \]

and the contravariant version is

\[ (\nu^\alpha) = \mp v^{-1} e^{-\psi}(1, u^i). \]

Thus, we have

1.1. Remark. Let \( M \) be spacelike graph in a future oriented coordinate system. Then the contravariant future directed normal vector has the form

\[ (\nu^\alpha) = v^{-1} e^{-\psi}(1, u^i) \]

and the past directed

\[ (\nu^\alpha) = -v^{-1} e^{-\psi}(1, u^i). \]
In the Gauß formula (1.2) we are free to choose the future or past directed normal, but we stipulate that we always use the past directed normal for reasons that we have explained in [5, Section 2].

Look at the component $\alpha = 0$ in (1.2) and obtain in view of (1.15)

$$e^{-\psi}v^{-1}h_{ij} = -u_{ij} - \bar{\Gamma}_{00}^0u_iu_j - \bar{\Gamma}_{0j}^0u_i - \bar{\Gamma}_{0i}^0u_j - \bar{\Gamma}_{ij}^0.$$  

Here, the covariant derivatives are taken with respect to the induced metric of $M$, and

$$(1.17) -\bar{\Gamma}_{ij}^0 = e^{-\psi}\bar{h}_{ij},$$

where $\bar{h}_{ij}$ is the second fundamental form of the hypersurfaces $\{x^0 = \text{const}\}$.

An easy calculation shows

$$\bar{h}_{ij}e^{-\psi} = -\frac{1}{2}\dot{\sigma}_{ij} - \psi\sigma_{ij},$$

where the dot indicates differentiation with respect to $x^0$.

Next we shall define the strong volume decay condition.

1.2. Definition. Suppose there exists a time function $x^0$ such that the future end of $N$ is determined by $\{\tau_0 \leq x^0 < b\}$ and the coordinate slices $M_\tau = \{x^0 = \tau\}$ have positive mean curvature with respect to the past directed normal for $\tau_0 \leq \tau < b$. In addition the volume $|M_\tau|$ should satisfy

$$\lim_{\tau \to b} |M_\tau| = 0.$$  

A decay like that is normally associated with a future singularity and we simply call it volume decay. If $(g_{ij})$ is the induced metric of $M_\tau$ and $g = \det(g_{ij})$, then we have

$$(1.20) \log g(\tau_0, x) - \log g(\tau, x) = \int_{\tau_0}^{\tau} e^{\psi}\bar{H}(s, x) \quad \forall x \in S_0,$$

where $\bar{H}(\tau, x)$ is the mean curvature of $M_\tau$ in $(\tau, x)$. For a proof we refer to [7].

In view of (1.19) the left-hand side of this equation tends to infinity if $\tau$ approaches $b$ for a.e. $x \in S_0$, i.e.,

$$\lim_{\tau \to b} \int_{\tau_0}^{\tau} e^{\psi}\bar{H}(s, x) = \infty \quad \text{for a.e. } x \in S_0.$$  

Assume now, there exists a continuous, positive function $\varphi = \varphi(\tau)$ such that

$$(1.22) e^{\psi}\bar{H}(\tau, x) \geq \varphi(\tau) \quad \forall (\tau, x) \in (\tau_0, b) \times S_0,$$

where

$$(1.23) \int_{\tau_0}^{b} \varphi(\tau) = \infty,$$

then we say that the future of $N$ satisfies a strong volume decay condition.
1.3. Remark. (i) By approximation we may—and shall—assume that the function $\varphi$ above is smooth.

(ii) A similar definition holds for the past of $N$ by simply reversing the time direction. Notice that in this case the mean curvature of the coordinate slices has to be negative.

1.4. Lemma. Suppose that the future of $N$ satisfies a strong volume decay condition, then there exist a time function $\tilde{x}^0 = \tilde{x}^0(x^0)$, where $x^0$ is the time function in the strong volume decay condition, such that the mean curvature $\bar{H}$ of the slices $\tilde{x}^0 = \text{const}$ satisfies the estimate

$$e^{\tilde{\psi}} \bar{H} \geq 1.$$  

(1.24)

The factor $e^{\tilde{\psi}}$ is now the conformal factor in the representation

$$ds^2 = e^{2\tilde{\psi}} (-d\tilde{x}^0)^2 + \sigma_{ij}dx^i dx^j.$$  

(1.25)

The range of $\tilde{x}^0$ is equal to the interval $[0, \infty)$, i.e., the singularity corresponds to $\tilde{x}^0 = \infty$.

Proof. Define $\tilde{x}^0$ by

$$\tilde{x}^0 = \int_{\tau_0}^{x^0} \varphi(\tau),$$  

(1.26)

where $\varphi$ is the function in (1.22) now assumed to be smooth.

The conformal factor in (1.25) is then equal to

$$e^{2\tilde{\psi}} = e^{2\psi} \frac{\partial x^0}{\partial \tilde{x}^0} \frac{\partial x^0}{\partial \tilde{x}^0} = e^{2\psi} \varphi^{-2},$$  

(1.27)

and hence

$$e^{\tilde{\psi}} \bar{H} = e^{\psi} \bar{H} \varphi^{-1} \geq 1,$$  

(1.28)

in view of (1.22). □

As we mentioned in the introduction there are spacetimes which satisfy a mean curvature barrier condition but the resulting singularity is not crushing.

To construct an example let us start with a $S$-AdS$_{(n+2)}$ spacetime with metric

$$ds^2 = -f dt^2 + f^{-1} dr^2 + r^2 \sigma_{ij} dx^i dx^j,$$  

(1.29)

where

$$f = \kappa - \frac{2}{n(n+1)} A r^2 - m r^{-(n-1)}$$  

(1.30)

with constants $A$ and $m > 0$; $(\sigma_{ij})$ is the metric of a compact $n$-dimensional spaceform of curvature $\kappa = 0, 1, -1$.

This spacetime satisfies the Einstein equations

$$G_{\alpha\beta} + A \tilde{g}_{\alpha\beta} = 0.$$  

(1.31)
Let us suppose for simplicity that $\kappa = 1$ and $\Lambda < 0$, though this is not important in our considerations. In $\{ r = 0 \}$ is a black hole singularity and the event horizon $\mathcal{H} = f^{-1}(0)$ is characterized by $r = r_0$.

The region $\{ f < 0 \}$ is the black hole region. In this region $r$ is the time function and $t$ is a spatial variable. Let us pick the black hole region. Normally the variable $t$ describes the real axis, but, since it is a spatial variable, we are free to compactify it, and we shall suppose that $t$ is a variable for $S^1$. By this compactification we have defined a globally hyperbolic spacetime $N$ with compact Cauchy hypersurface $S_0 = S^1 \times S^n$ which satisfies the timelike convergence condition since

$$R_{\alpha\beta} = \frac{2}{n} \Lambda g_{\alpha\beta}$$

and $\Lambda$ is supposed to be negative.

$N$ has a crushing singularity in $r = 0$, and, as we shall show in a moment, also a mean curvature barrier singularity in $r = r_0$, which is however not crushing, since the metric quantities were not changed by the compactification but only the topology.

Define

$$f = -f \quad \text{and} \quad \psi = -\frac{1}{2} \log \tilde{f},$$

then the metric can be expressed as

$$ds^2 = e^{2\psi} (-dr^2 + \tilde{f}^2 dt^2 + \tilde{f}^2 \sigma_{ij} dx^i dx^j)$$

$$\equiv e^{2\psi} (-dr^2 + \tilde{\sigma}_{ab} dx^a dx^b).$$

The second fundamental form of the hypersurfaces $\{ r = \text{const} \}$ with respect to the past directed normal is given by

$$e^{-\psi} h_{ab} = \frac{1}{2} \tilde{\sigma}_{ab} - \frac{1}{2} \tilde{f}^{-1} \tilde{f}_{ab},$$

where the dot indicates differentiation with respect to $r$, and where we note that the time function $r$ is past directed in contrast to the usual convention. Hence the mean curvature $\bar{H}$ is equal to

$$\bar{H} = \tilde{f}^{-\frac{1}{2}} (\tilde{f}^{-1} + n \tilde{f} r^{-1})$$

and we deduce that $\bar{H}$ tends to $-\infty$, if the hypersurfaces approach the horizon $\mathcal{H}$, and to $\infty$, if the hypersurfaces approach the black hole singularity $r = 0$.

Sometimes, we need a Riemannian reference metric, e.g., if we want to estimate tensors. Since the Lorentzian metric can be expressed as

$$\bar{g}_{\alpha\beta} dx^\alpha dx^\beta = e^{2\psi} (-dx^0^2 + \sigma_{ij} dx^i dx^j),$$

we define a Riemannian reference metric ($\tilde{g}_{\alpha\beta}$) by

$$\tilde{g}_{\alpha\beta} dx^\alpha dx^\beta = e^{2\psi} (dx^0^2 + \sigma_{ij} dx^i dx^j)$$

and we abbreviate the corresponding norm of a vectorfield $\eta$ by

$$||\eta|| = (\tilde{g}_{\alpha\beta} \eta^\alpha \eta^\beta)^{1/2}.$$
with similar notations for higher order tensors.

2. The evolution problem

The evolution problem (0.2) is a parabolic problem, hence a solution exists on a maximal time interval \([0, T^*)\), \(0 < T^* \leq \infty\), cf. [4, Section 2], where we apologize for the ambiguity of also calling the evolution parameter \(t\).

Next, we want to show how the metric, the second fundamental form, and the normal vector of the hypersurfaces \(M(t)\) evolve. All time derivatives are total derivatives. We refer to [5] for more general results and to [4, Section 3], where proofs are given in a Riemannian setting, but these proofs are also valid in a Lorentzian environment.

2.1. Lemma. The metric, the normal vector, and the second fundamental form of \(M(t)\) satisfy the evolution equations

\[
\begin{align*}
\dot{g}_{ij} &= -2H^{-1}h_{ij}, \\
\dot{\nu} &= \nabla_M(-H^{-1}) = g^{ij}(-H^{-1})_i x_j,
\end{align*}
\]

and

\[
\begin{align*}
\dot{h}^i_j &= (-H^{-1})^i_j + H^{-1}h^k_i h^j_k + H^{-1}\bar{R}_{\alpha\beta\gamma\delta} x^\alpha_i x^\beta_j x^\gamma_k x^\delta_l h^k_l, \\
\dot{h}_{ij} &= (-H^{-1})_{ij} - H^{-1}h^k_i h_{kj} + H^{-1}\bar{R}_{\alpha\beta\gamma\delta} x^\alpha_i x^\beta_j x^\gamma_k x^\delta_l.
\end{align*}
\]

2.2. Lemma (Evolution of \(H^{-1}\)). The term \(H^{-1}\) evolves according to the equation

\[
(H^{-1})' - H^{-2} \Delta H^{-1} = -H^{-2}(\|A\|^2 + \bar{R}_\alpha\beta\gamma\delta x^\alpha_i x^\beta_j x^\gamma_k x^\delta_l)H^{-1}
\]

where

\[
(H^{-1})' = \frac{d}{dt}H^{-1}
\]

and

\[
\|A\|^2 = h_{ij} h^{ij}.
\]

From (2.3) we deduce with the help of the Ricci identities and the Codazzi equations a parabolic equation for the second fundamental form.
2.3. Lemma. The mixed tensor $h_{ij}$ satisfies the parabolic equation

$$
\dot{h}_{ij} - H^{-2}\Delta h_{ij} = -H^{-2}\|A\|^2 h_{ij} + 2H^{-1}h_{ik}h_{kj} - 2H^{-3}H_{ij} + 2H^{-2}\bar{\Omega}_{\alpha\beta\gamma\delta} x_{\alpha}^m x_{\beta}^i x_{\gamma}^k x_{\delta}^j h_{km} g^{rij}
$$

(2.8)

Since the timelike convergence condition is assumed to be valid we immediately deduce from Lemma 2.2

2.4. Lemma. There exists a positive constant $c_0 = c_0(M_0)$, such that the estimate

$$
H \geq c_0 e^{\frac{1}{4}t}
$$

is valid during the evolution.

Proof. Let $\varphi = H^{-1}e^{\frac{1}{2}t}$, then $\varphi$ satisfies the inequality

$$
\dot{\varphi} - H^{-2}\Delta \varphi \leq -H^{-2}\|A\|^2 \varphi + \frac{1}{n} \varphi \leq 0,
$$

(2.10)

hence we conclude

$$
\varphi \leq \sup_{M_0} \varphi = \sup_{M_0} H.
$$

(2.11)

3. LOWER ORDER ESTIMATES

The evolution problem (0.2) exists on a maximal time interval $I = [0, T^*)$. We want to prove that $T^* = \infty$, and that the flow hypersurfaces $M(t)$ run into the future singularity, if $t$ tends to infinity.

The latter property is a characteristic of the inverse mean curvature flow under very weak assumptions: if the flow exists for all time, then it cannot stay in a compact region of $N$, or, more precisely

3.1. Lemma. Let $N$ be a cosmological spacetime with a future mean curvature barrier, and let $M_0$ be a compact spacelike hypersurface with positive mean curvature. Suppose that $N = \mathbb{R} \times S_0$ and that the metric is given as in (0.4). Assume that the inverse mean curvature flow with initial hypersurface $M_0$ exists for all time, and let the flow hypersurfaces $M(t)$ be expressed as graphs of a function $u$ over $S_0$

$$
M(t) = \{(x^0, x): x^0 = u(t, x), x \in S_0\}.
$$

(3.1)
Then there holds

\[(3.2) \quad \lim_{t \to \infty} \inf_{S_0} u(t, \cdot) = \infty.\]

**Proof.** (i) Because of the barrier condition a future end of $N$, $N_+$, can be foliated by hypersurfaces of positive constant curvature and we can choose the mean curvature $\tau$ of that CMC foliation as new time function $x^0 = \tau$ in $N_+$.

\[(3.3) \quad N_+ = \{ (\tau, x) : k \leq \tau < \infty, x \in S_0 \},\]

cf. Remark 0.2, where $k$ is a positive constant and where we used the same symbol $S_0$ for the compact Cauchy hypersurface—indeed, we could use the original Cauchy hypersurface $S_0$, since it need not be a level hypersurface.

Let $t_0$ be such that

\[(3.4) \quad c_0 e^{\frac{t}{c_0}} > 2k,\]

where $c_0$ is the constant in inequality (2.9), then we claim that

\[(3.5) \quad M(t) \subset N_+ \quad \forall t \geq t_0.\]

To prove this claim we shall apply the Synge’s lemma. Denote the coordinate slices $x^0 = \tau$ by $M_\tau$, i.e., $M_\tau$ has constant mean curvature $\bar{H} = \tau$.

It suffices to show that all $M(t)$ with $t \geq t_0$ lie in the future of $M_k$. Suppose this were not the case for some $M(t)$, then the Lorentzian distance between $M(t)$ and $M_k$ would be positive

\[(3.6) \quad d = d(M(t), M_k) > 0\]

and hence there would exist a maximal future directed geodesic $\gamma$ from $M(t)$ to $M_k$. Synge’s lemma would then yield

\[(3.7) \quad H_{|M_t}(\gamma(d)) \geq H_{|M(t)}(\gamma(0)) + \int_0^d \bar{R}_{\alpha\beta} \gamma^\alpha \gamma^\beta;\]

a contradiction in view of (3.4) and the timelike convergence condition.

(ii) Thus, the flow hypersurfaces $M(t)$ are covered by the new coordinate system for $t \geq t_0$. The metric of $N$ has again the form as in (0.4).

Now, the mean curvature $\bar{H}$ of the coordinate slices satisfies the evolution equation

\[(3.8) \quad \dot{\bar{H}} = -\Delta e^\psi + (|\bar{A}|^2 + \bar{R}_{\alpha\beta} \nu^\alpha \nu^\beta) e^\psi,\]

where the dot indicates differentiation with respect to $x^0$, the Laplace operator is the Laplace Beltrami operator of the slice, $|\bar{A}|^2$ the square of the second fundamental form and $\nu$ the past directed normal and $e^\psi$ the conformal factor of the metric.
This relation is valid for the slices of any time function $x^0$ for which the metric has the form as in (0.4), since the slices are solutions of the evolution equation
\begin{equation}
\dot{x} = -e^\psi \nu
\end{equation}
from which the relation (3.8) can be easily deduced: apply the general formula (3.8) in [5] with $\Phi = -e^\psi$.

For the special time function $x^0 = \tau$ we therefore obtain
\begin{equation}
1 = \dot{H} \geq -\Delta e^\psi + \frac{1}{n} \tau^2 e^\psi.
\end{equation}

Moreover, let $x_0 \in S_0$ be a point where, for fixed $\tau$,
\begin{equation}
sup_{S_0} e^\psi(\tau, \cdot) = e^\psi(\tau, x_0),
\end{equation}
then the maximum principle implies
\begin{equation}
1 \geq \frac{1}{n} \tau^2 e^\psi(\tau, x_0) \geq \frac{1}{n} \tau^2 e^\psi(\tau, x) \quad \forall x \in S_0
\end{equation}
and hence
\begin{equation}
\dot{H} e^\psi \leq n\dot{H}^{-1}
\end{equation}
for all slices $M_\tau$.

This inequality will be the key ingredient to prove the limit relation (3.2).

(iii) Define the function $\varphi$ on $t \geq t_0$ by
\begin{equation}
\varphi(t) = \inf_{S_0} u(t, \cdot),
\end{equation}
then $\varphi$ is Lipschitz continuous and for a.e. $t$ there holds
\begin{equation}
\dot{\varphi}(t) = \dot{u}(t, x_t),
\end{equation}
where $x_t$ is such that the infimum is attained in $x_t$. This result is well known; we shall give a short prove in Lemma 3.2 below for the sake of completeness.

Now, from (0.2), looking at the component $\alpha = 0$, we deduce that $u$ satisfies the evolution equation
\begin{equation}
\dot{u} = \frac{\tilde{v}}{H e^\psi},
\end{equation}
where $\tilde{v} = v^{-1}$ and where the time derivative is the total derivative, i.e.,
\begin{equation}
\dot{u} = \frac{\partial u}{\partial t} + u_i \dot{x}^i
\end{equation}
and hence
\begin{equation}
\frac{\partial u}{\partial t} = \frac{v}{H e^\psi}.
\end{equation}

From (1.16) we infer
\begin{equation}
e^{-\psi} \tilde{v} H = -\Delta u - \tilde{f}_0^0 \|Du\|^2 - 2\tilde{f}_0^0 u^i + e^{-\psi} \dot{H},
\end{equation}
and conclude further, with the help of the maximum principle, that in $x_t$
\begin{equation}
H \leq \dot{H},
\end{equation}
and thus
\[(3.21) \quad \frac{\partial u}{\partial t} \geq \frac{1}{\bar{H}e^\psi}\]
in \(x_t\).
Therefore, \(\varphi\) satisfies
\[(3.22) \quad \dot{\varphi} \geq \frac{1}{\bar{H}e^\psi} \quad \text{for a.e. } t \geq t_0,\]
hence
\[(3.23) \quad \dot{\varphi} \geq \frac{1}{n} \bar{H} = \frac{1}{n} \varphi\]
in view of (3.13) and the fact that the slices \(M_\tau\) have mean curvature \(\tau\).
From this inequality we immediately deduce
\[(3.24) \quad \varphi(t) \geq \varphi(t_0)e^{\frac{1}{n}(t-t_0)} \quad \forall t \geq t_0\]
proving the lemma. \(\square\)

3.2. Lemma. Let \(S_0\) be compact and \(f \in C^1(J \times S_0)\), where \(J\) is any open interval, then
\[(3.25) \quad \varphi(t) = \inf_{S_0} f(t, \cdot)\]
is Lipschitz continuous and there holds a.e.
\[(3.26) \quad \dot{\varphi} = \frac{\partial f}{\partial t}(t, x_t),\]
where \(x_t\) is a point in which the infimum is attained.
A corresponding result is also valid if \(\varphi\) is defined by taking the supremum instead of the infimum.

Proof. \(\varphi\) is obviously Lipschitz continuous and thus a.e. differentiable by Rademacher’s theorem.
For arbitrary \(t_1, t_2 \in J\) we have
\[(3.27) \quad \varphi(t_1) - \varphi(t_2) = f(t_1, x_{t_1}) - f(t_2, x_{t_2}) = f(t_1, x_{t_1}) - f(t_2, x_{t_1}) \geq f(t_1, x_{t_1}) - f(t_2, x_{t_1}).\]
Now, let \(\varphi\) be differentiable at \(t_1\), then, by choosing \(t_2 > t_1\), and looking at the difference quotients of both sides, we conclude
\[(3.28) \quad \dot{\varphi}(t_1) \leq \frac{\partial f}{\partial t}(t_1, x_{t_1}).\]
Choosing \(t_2 < t_1\) we obtain the opposite inequality, completing the proof of the lemma. \(\square\)

We have proved that the flow hypersurfaces run straight in the singularity, if the flow exists for all time. However, it might happen that the flow runs into the future singularity in finite time.
To exclude this possibility we have imposed the strong volume decay condition
3.3. Lemma. Let $N$ satisfy a strong volume decay condition with respect to the future, then, for any finite $T$, $0 < T \leq T^*$, the flow stays in a precompact region $\Omega_T$ for $0 \leq t < T$.

Proof. According to Lemma 1.4 we may choose a time function $x^0$ such that the relation (1.24) is valid for the coordinate slices $x^0 = \text{const}$.

Let $M(t) = \text{graph } u$ be the flow hypersurfaces, and set

$$\varphi(t) = \sup_{S_0} u(t, \cdot).$$

Then, similarly as in the proof of Lemma 3.1 we deduce that for a.e. $t$

$$\dot{\varphi} = \frac{1}{He^\psi} \leq \frac{1}{\bar{H}e^\psi} \leq 1,$$

in view of (1.24).

Hence we infer

$$\varphi \leq \varphi(0) + t \quad \forall 0 \leq t < T^*,$$

which proves the lemma since the singularity corresponds to $x^0 = \infty$. \hfill \Box

4. $C^1$-estimates

We consider a smooth solution of the evolution equation (0.2) in a maximal time interval $[0, T^*)$ and shall prove a priori estimates for

$$\tilde{v} = v^{-1} = \frac{1}{\sqrt{1 - |Du|^2}}$$

in $Q_T = [0, T] \times S_0$ for any $0 < T < T^*$.

The proof is a slight modification of the proof of the corresponding result for the mean curvature flow in [6]. We note that the timelike convergence condition is not necessary for this estimate.

Let us first state an evolution equation for $\tilde{v}$.

4.1. Lemma (Evolution of $\tilde{v}$). The quantity $\tilde{v}$ satisfies the evolution equation

$$\dot{\tilde{v}} - H^{-2} \Delta \tilde{v} = -H^{-2} A^2 \tilde{v} - 2H^{-1} \eta_{\alpha \beta} \nu^\alpha \nu^\beta$$

$$-2H^{-2} h^{ij} x_i^\alpha x_j^\beta \eta_{\alpha \beta} - H^{-2} g^{ij} \eta_{\alpha \beta} x_i^\alpha x_j^\beta \nu^\alpha$$

$$- H^{-2} \bar{R} \eta_{\alpha \beta} x_k^\beta \eta_{\gamma} x_l^{\gamma} g^{kl},$$

where $\eta$ is the covariant vector field $(\eta_\alpha) = e^{\psi}(-1, 0, \ldots, 0)$.

Proof. We have $\tilde{v} = \langle \eta, \nu \rangle$. Let $(\xi^i)$ be local coordinates for $M(t)$. Differentiating $\tilde{v}$ yields the result, cf. [6] Lemma 3.2 for details. \hfill \Box
4.2. Lemma. Consider the flow in a precompact region $\Omega$, then there exists a constant $c = c(\Omega)$ such that for any positive function $0 < \epsilon = \epsilon(x)$ on $S_0$ and any hypersurface $M(t) \subset \Omega$ of the flow we have
\begin{align}
\|\nu\| &\leq c\tilde{v}, \\
g^{ij} &\leq c\tilde{v}^2\sigma^{ij},
\end{align}
and
\begin{equation}
|h^{ij}\eta_\alpha x^\alpha_i x^\beta_j| \leq \frac{\epsilon}{2}\|A\|^2\tilde{v} + \frac{c}{2\epsilon}\tilde{v}^3
\end{equation}
where $(\eta_\alpha)$ is the vector field in Lemma 4.1.

Confer [6, Lemma 3.3] for a proof.

Combining the preceding lemmata we infer

4.3. Lemma. Consider the flow in a precompact region $\Omega$, then there exists a constant $c = c(\Omega)$ such that for any positive function $\epsilon = \epsilon(x)$ on $S_0$ the term $\tilde{v}$ satisfies a parabolic inequality of the form
\begin{equation}
\dot{\tilde{v}} - H^{-2}\Delta\tilde{v} \leq -(1 - \epsilon)H^{-2}\|A\|^2\tilde{v} + cH^{-2}[1 + \epsilon^{-1}]\tilde{v}^3.
\end{equation}

Proof. The terms on the right-hand side of (4.2) having a factor $H^{-2}$ can obviously be estimated as claimed.

The remaining term can be estimated by
\begin{equation}
2H^{-1}|\eta_\alpha\nu^\alpha\nu^\beta| \leq 2cH^{-1}\tilde{v}^2 \\
\leq \frac{\epsilon}{2}\tilde{v} + 2nc\epsilon^{-1}H^{-2}\tilde{v}^3.
\end{equation}
The claim then follows from the relation
\begin{equation}
\frac{1}{n}H^2 \leq |A|^2,
\end{equation}
i.e.,
\begin{equation}
-H^{-2}|A|^2\tilde{v} \leq -\frac{1}{n}\tilde{v}.
\end{equation}

□

We further need the following two lemmata

4.4. Lemma. Let $M(t) = \text{graph} u(t)$ be the flow hypersurfaces, then we have
\begin{align}
\dot{u} - H^{-2}\Delta u &\geq 2e^{-\psi}\tilde{v}H^{-1} - H^{-2}e^{-\psi}g^{ij}h_{ij} \\
&\quad + H^{-2}f^0_{0i}\|Du\| + 2H^{-2}f^0_{0i}u^i,
\end{align}
where the time derivative is a total derivative.

Proof. We use the relation
\begin{equation}
\dot{u} = e^{-\psi}\tilde{v}H^{-1}
\end{equation}
and
\begin{equation}
\text{[1.16]},
\end{equation}
\end{equation}
4.5. Lemma. Let \( \Omega \subset N \) be precompact and \( M \subset \Omega \) be a spacelike graph over \( S_0 \), \( M = \text{graph } u \), then
\[
|\tilde{v}_i u^i| \leq c\tilde{v}^3 + \|A\|e^\psi \|Du\|^2,
\]
where \( c = c(\Omega) \).

Proof. Confer the proof of [6, Lemma 3.6]. \( \square \)

We are now ready to prove the a priori estimate for \( \tilde{v} \).

4.6. Lemma. Let \( \Omega \subset N \) be precompact. Then, as long as the flow stays in \( \Omega \), the term \( \tilde{v} \) is a priori bounded
\[
\tilde{v} \leq c = c(\Omega, \sup_{M_0} \tilde{v}).
\]
In particular, we do not have to assume that the timelike convergence is valid, and we note that \( c \) does not depend explicitly on \( T \).

Proof. Let \( \mu, \lambda \) be positive constants, where \( \mu \) is supposed to be small and \( \lambda \) large, and define
\[
\phi = e^{\mu e^{-\lambda u}},
\]
where we assume without loss of generality that \( u \leq -1 \), otherwise replace in (4.14) \( u \) by \( (u - c) \), \( c \) large enough.

We shall show that
\[
w = \tilde{v}\phi
\]
is a priori bounded as indicated in (4.13) if \( \mu, \lambda \) are chosen appropriately.

In view of Lemma 4.2 and Lemma 4.4 we have
\[
\dot{\phi} - H^{-2}\Delta \phi \leq c\mu\lambda e^{-\lambda u}H^{-2}\tilde{v}^2\phi - \mu\lambda^2 e^{-\lambda u}[1 + \mu e^{-\lambda u}]H^{-2}\|Du\|^2\phi,
\]
since \( 0 < H \), from which we further deduce, taking Lemma 4.3 and Lemma 4.5 into account,
\[
\dot{w} - H^{-2}\Delta w \leq -(1 - \epsilon)H^{-2}\|A\|^2\tilde{v}\phi + cH^{-2}[1 + \epsilon^{-1}]\tilde{v}^3\phi
\]
\[
- \mu\lambda^2 e^{-\lambda u}[1 + \mu e^{-\lambda u}]H^{-2}\tilde{v}\|Du\|^2\phi
\]
\[
+ c\mu\lambda e^{-\lambda u}H^{-2}\tilde{v}^3\phi + 2\mu\lambda e^{-\lambda u}H^{-2}\|A\|e^\psi \|Du\|^2\phi.
\]

We estimate the last term on the right-hand side by
\[
2\mu\lambda e^{-\lambda u}H^{-2}\|A\|e^\psi \|Du\|^2\phi \leq (1 - \epsilon)H^{-2}\|A\|^2\tilde{v}\phi
\]
\[
+ \frac{1}{1 - \epsilon} \mu^2\lambda^2 e^{-2\lambda u}H^{-2}\tilde{v}^{-1}\epsilon^2\psi \|Du\|^4\phi.
\]
and conclude
\[ w - H^{-2} \Delta w \leq c[1 + \epsilon^{-1}]H^{-2}\tilde{v}^3 \varphi \]
(4.19)
\[ + \left[ \frac{1}{1-\epsilon} - 1 \right] \mu^2 \lambda^2 e^{-2\lambda u} H^{-2} \| Du \|^2 \tilde{v} \varphi \]
\[ - \mu \lambda^2 e^{-\lambda u} H^{-2} \| Du \|^2 \tilde{v} \varphi, \]
where we have used that
(4.20)
\[ e^{2\psi} \| Du \|^2 \leq \tilde{v} \]

Setting \( \epsilon = e^{\lambda u} \), we then obtain
(4.21)
\[ H^2 (w - H^{-2} \Delta w) \leq c e^{-\lambda u} \tilde{v}^3 \varphi + c \mu \lambda e^{-\lambda u} \tilde{v}^3 \varphi \]
\[ + \left[ \frac{\mu}{1-\epsilon} - 1 \right] \mu \lambda^2 e^{-\lambda u} \| Du \|^2 \tilde{v} \varphi. \]

Now, we choose \( \mu = \frac{1}{2} \) and \( \lambda_0 \) so large that
(4.22)
\[ \frac{\mu}{1 - e^{\lambda u}} \leq \frac{3}{4} \quad \forall \lambda \geq \lambda_0, \]
and infer that the last term on the right-hand side of (4.21) is less than
(4.23)
\[ -\frac{1}{8} \lambda^2 e^{-\lambda u} \| Du \|^2 \tilde{v} \varphi \]
which in turn can be estimated from above by
(4.24)
\[ -c \lambda^2 e^{-\lambda u} \tilde{v}^3 \varphi \]
at points where \( \tilde{v} \geq 2 \).

Thus, we conclude that for
(4.25)
\[ \lambda \geq \max(\lambda_0, 4) \]
the parabolic maximum principle, applied to \( w \), yields
(4.26)
\[ w \leq \text{const}(\| w(0) \|_{s_0}, \lambda_0, \Omega). \]
\[ \square \]

5. \( C^2 \)-estimates

We want to prove that, as long as the flow stays in a precompact set \( \Omega \subset N \), the principal curvatures of the flow hypersurfaces are a priori bounded by a constant depending only on \( \Omega \) and the initial hypersurface \( M_0 \). Again we do not need the timelike convergence condition for this estimate.

Let us first prove an a priori estimate for \( H \).

5.1. Lemma. Let \( \Omega \subset N \) be precompact and assume that the flow (0.2) stays in \( \Omega \) for \( 0 \leq t \leq T < T^* \), then the mean curvature of the flow hypersurfaces is bounded by
(5.1)
\[ 0 < H \leq c(\Omega, \sup M_0 H). \]
Proof. From Lemma 2.2 we immediately deduce that \( \varphi = \log H \) satisfies the evolution equation
\[
\dot{\varphi} - H^{-2} \Delta \varphi = H^{-2} (|A|^2 + \bar{R}_{\alpha \beta} \nu^\alpha \nu^\beta) - H^{-2} \| D \varphi \|^2.
\]

Let \( \lambda \) be large and set
\[
w = \varphi + \lambda \hat{\nu}.
\]
Then we conclude from (4.6) that \( w \) satisfies the parabolic inequality
\[
\dot{w} - H^{-2} \Delta w \leq - \frac{\lambda}{2} H^{-2} |A|^2 + c \lambda H^{-2},
\]
if \( \lambda \) is large enough, \( \lambda \geq \lambda(\Omega) \). Hence the parabolic maximum principle yields the result in view of the relation
\[
\frac{1}{\lambda} H^2 \leq |A|^2.
\]

5.2. Lemma. Under the assumptions of Lemma 5.1 the principal curvatures \( \kappa_i, 1 \leq i \leq n \), of the flow hypersurfaces are a priori bounded in \( \Omega \)
\[
|\kappa_i| \leq c(\Omega, \sup_{M_0} |A|).
\]

Proof. Since \( 0 \leq H \), it suffices to estimate
\[
\sup_i \kappa_i \leq c(\Omega, \sup_{M_0} |A|).
\]

Let \( \varphi \) be defined by
\[
\varphi = \sup \{ h_{ij} \eta^i \eta^j : \| \eta \| = 1 \}.
\]
We claim that \( \varphi \) is a priori bounded in \( \Omega \).

Let \( 0 < T < T^* \), and \( x_0 = x_0(t_0) \), with \( 0 < t_0 \leq T \), be a point in \( M(t_0) \) such that
\[
\sup_{M_0} \varphi < \sup_{M(t)} \{ \sup_{M(t)} \varphi : 0 < t \leq T \} = \varphi(x_0).
\]

We then introduce a Riemannian normal coordinate system \( (\xi^i) \) at \( x = x(t_0, \xi_0) \) we have
\[
g_{ij} = \delta_{ij} \quad \text{and} \quad \varphi = h^n_\nu.
\]

Let \( \tilde{\eta} = (\tilde{\eta}^i) \) be the contravariant vector field defined by
\[
\tilde{\eta} = (0, \ldots, 0, 1),
\]
and set
\[
\tilde{\varphi} = h_{ij} \tilde{\eta}^i \tilde{\eta}^j / g_{ij} \tilde{\eta}^i \tilde{\eta}^j.
\]

\( \tilde{\varphi} \) is well defined in neighbourhood of \( (t_0, \xi_0) \), and \( \tilde{\varphi} \) assumes its maximum at \( (t_0, \xi_0) \). Moreover, at \( (t_0, \xi_0) \) we have
\[
\dot{\tilde{\varphi}} = \hat{h}^n_\nu,
\]
and the spatial derivatives do also coincide; in short, at \((t_0, \xi_0)\) \(\tilde{\varphi}\) satisfies the same differential equation 2.8 as \(h^*_n\). For the sake of greater clarity, let us therefore treat \(h^*_n\) like a scalar and pretend that \(\varphi = h^*_n\).

At \((t_0, \xi_0)\) we have \(\dot{\varphi} \geq 0\), and, in view of the maximum principle, we deduce from Lemma 2.3

\[
0 \leq H^{-2}(\|A\|^2 h^*_n + c|h^*_n|^2 + c).
\]

Thus \(\varphi\) is apriori bounded in \(\Omega\) by a constant \(c\) depending only on \(\Omega\) and the initial hypersurface \(M_0\).

\[\square\]

6. Longtime existence

Let us look at the scalar version of the flow as in (3.18)

\[
\frac{\partial u}{\partial t} = e^{-\psi} v H^{-1}
\]

defined in the cylinder

\[
Q_{T^*} = [0, T^*) \times S_0
\]

with initial value \(u(0) \in C^\infty(S_0)\).

Suppose that \(T^* < \infty\), then, from Lemma 3.3 we conclude that the flow stays in a compact region of \(N\). Furthermore, in view of Lemma 4.6 and the \(C^2\)-estimates of Section 5 we obtain uniform \(C^2\)-estimates for \(u\).

Thus, the differential operator on the right-hand side of (6.1) is uniformly elliptic in \(u\) independent of \(t\), since there are constants \(c_1, c_2\) such that

\[
0 < c_1 \leq H \leq c_2 \quad \forall 0 \leq t < T^*,
\]

in view of Lemma 2.4.

Hence, we can apply the known regularity results, cf. e.g., [13, Chap. 5.5], to conclude that uniform \(C^{2,\alpha}\)-estimates are valid, leading further to uniform \(C^{m,\alpha}\)-estimates for any \(m \in \mathbb{N}\), due to the regularity result for linear operators. But this will contradict the maximality of \(T^*\).

Therefore, \(T^* = \infty\), i.e., the flow exists for all time, and for any finite \(T\) we have a priori estimates in \(C^m([0, T] \times S_0)\) for any \(m \in \mathbb{N}\).

7. A new time function

We know that the flow exists for all time and hence we conclude from Lemma 3.1 and Lemma 3.3 that the flow hypersurfaces provide a foliation of the future of \(M_0\), i.e., the flow parameter \(t\) could be used as a new time function in \(D^+(M_0)\), if \(Dt\) is timelike.

7.1. Lemma. The flow parameter \(t\) can be used as future directed time function in \(D^+(M_0)\).
Proof. Let \((x^\alpha)\) be a future directed coordinate system such that the relation (0.4) is valid. Then look at the scalar version of the flow, equation (6.1). If we can show that \((\tilde{x}^\alpha)\) with

\[
\tilde{x}^0 = t, \quad \tilde{x}^i = x^i
\]

represents a regular coordinate transformation with positive Jacobi determinant, then the lemma is proved.

Now, the inverse coordinate transformation \(x = x(\tilde{x})\), which exists, since we already know that the flow hypersurfaces provide a foliation, has the form

\[
x^0 = u(t, x) \equiv u(\tilde{x}), \quad x^i = \tilde{x}^i,
\]

where we apologize for using the same symbol \(x\) to represent an \((n+1)\)-tupel as well as the space coordinates \((x^i)\).

We immediately deduce

\[
\left| \frac{\partial x}{\partial \tilde{x}} \right| = \frac{\partial u}{\partial t} > 0,
\]

hence the result in view of the inverse function theorem. \(\square\)

The strong volume decay condition is not only sufficient to prove the long time existence of the inverse mean curvature flow, but also necessary.

7.2. Proposition. Let \(N\) be a cosmological spacetime, \(M_0 \subset N\) a compact, spacelike hypersurface with positive mean curvature, and suppose that the inverse mean curvature flow with initial hypersurface \(M_0\) exists for all time and provides a foliation of \(D^+(M_0)\), then \(N\) satisfies a future strong volume decay condition as well as a future mean curvature barrier condition.

Proof. Choose \(x^0 = t\) as new time function and let the metric of \(N\) be expressed as

\[
d\bar{s}^2 = e^{2\psi}(-dx^0)^2 + \sigma_{ij}(x^0, x)dx^i dx^j.
\]

\(M_0\) now replaces the Cauchy hypersurface \(S_0\) and the flow hypersurfaces \(M(t)\) are given as graphs of functions \(u\) with

\[
u(t, x) = t.
\]

Thus we conclude from (6.1) that

\[
1 = \frac{\partial u}{\partial t} = e^{-\psi}H^{-1},
\]

or equivalently,

\[
He^\psi = 1 \quad \forall x \in M(t),
\]

i.e., the strong volume decay condition is satisfied.

The mean curvature of the leaves \(M(t)\) tends to \(\infty\) in view of Lemma 2.4, hence \(N\) satisfies a future mean curvature barrier condition. \(\square\)
From now on, let us assume that $x^0 = t$ is the time function. Set
\begin{equation}
\tau = 1 - e^{-\frac{1}{\kappa} t},
\end{equation}
then the future spacetime singularity corresponds to $\tau = 1$, and there holds

7.3. Theorem. The quantity $1 - \tau$ can be looked at as the radius of the slices $\tau = \text{const}$ as well as a measure of the remaining life span of the spacetime, since we have
\begin{equation}
|M(\tau)| = |M_0|(1 - \tau)^n,
\end{equation}
and the length $L(\gamma)$ of any future directed curve starting from $M(\tau)$ is estimated from above by
\begin{equation}
L(\gamma) \leq c(1 - \tau),
\end{equation}
where
\begin{equation}
c = \frac{n}{\inf_{M_0} H}.
\end{equation}

Proof. Let $g = \det(g_{ij})$, where $(g_{ij})$ is the induced metric of $M(t) \equiv M(\tau)$, then
\begin{equation}
\frac{d}{dt}\sqrt{g} = -\sqrt{g}
\end{equation}
in view of (2.1), and hence
\begin{equation}
|M(\tau)| = |M_0|e^{-\frac{1}{\kappa} t} = |M_0|(1 - \tau)^n.
\end{equation}

To prove (7.10), we first note that in view of Lemma 2.4
\begin{equation}
H \geq \inf_{M_0} \frac{He^{\frac{1}{\kappa} t}}{M_0} = \frac{n}{\tau}(1 - \tau)^{-1},
\end{equation}
where $c$ is the constant in (7.11). One of Hawking’s singularity theorems then asserts that
\begin{equation}
L(\gamma) \leq c(1 - \tau),
\end{equation}
cf. [14, Prop. 37 on p. 288]. □

References
9. [Author], On the foliation of space-time by constant mean curvature hypersurfaces, 2003, e-print, 7 pages. [math.DG/0304122]


RUPRECHT-KARLS-UNIVERSITÄT, INSTITUT FÜR ANGEWANDE MATHematik, IM NEUENHEIMER FELD 294, 69120 HEIDELBERG, GERMANY
E-mail address: gerhardt@math.uni-heidelberg.de
URL: http://www.math.uni-heidelberg.de/studinfo/gerhardt/